

## ASSIGNMENT 7

### Reading:

105 Notes 8.1-8.3, 6.1-6.2 (again).  
Hand & Finch 4.7, 5.1.

1.

[This a (hopefully clearer) version of Hand & Finch 4.17, “tetherball”.] A mass  $m$  is attached to a weightless string that initially has a length  $s_0$ . The other end of the string is attached to a post of radius  $a$ . Neglect the effect of gravity. Suppose that the mass is set into motion. It is given an initial velocity of magnitude  $v_0$  directed so that the string remains taut. The string wraps itself around the post, causing the mass to spiral inward toward it.

(a)

Write the Lagrangian in terms of  $\dot{x}$  and  $\dot{y}$ , the cartesian velocity components of the mass. Is there a potential energy term?

(b)

Use as generalized coordinates  $s(t)$ , the length of the part of the string that is not yet in contact with the post, and  $\psi(t)$ , the azimuthal angle at which the string barely fails to make contact with the post. Express  $\dot{x}$  and  $\dot{y}$  in terms of these generalized coordinates and their time derivatives.

(c)

Write a (constraint) equation relating  $\dot{s}$  to  $\dot{\psi}$ . Use it to greatly simplify your answers for (b). Rewrite the Lagrangian using  $s$  as the only generalized coordinate.

(d)

Use the Euler-Lagrange equation to obtain an equation of motion for  $s$ . (You don't need to solve it.)

(e)

Since the Lagrangian has no explicit time dependence, and it depends quadratically on  $\dot{s}$ , the total energy is conserved. Write an equation setting the initial energy (expressed in terms of  $v_0$ ) equal to the energy at an arbitrary value of  $s$  (expressed in terms of  $s$  and  $\dot{s}$ ).

(f)

Use this equation to express  $dt$  in terms of  $ds$  multiplied by a function of  $s$ . Integrate it to solve for the time  $T$  that elapses before the mass hits the post. You should obtain the simple result

$$T = \frac{s_0^2}{2av_0}.$$

(g)

Is the angular momentum of the mass about the axis of the post conserved in this problem? Why or why not?

2.

Hand & Finch 4.19.

3.

Hand & Finch 4.21.

4.

Consider a particle of mass  $m$  that is constrained to move on the surface of a paraboloid whose equation (in cylindrical coordinates) is  $r^2 = 4az$ . If the particle is subject to a gravitational force  $-mg\hat{z}$ , show that the frequency of small oscillations about a circular orbit with radius  $\rho = \sqrt{4az_0}$  is

$$\omega = \sqrt{\frac{2g}{a + z_0}}.$$

5.

An orbit that is almost circular can be considered to be a circular orbit to which a small perturbation has been applied. Take  $\rho$  to be the (unperturbed) circular orbit radius and define

$$g(r) = \frac{1}{\mu} \frac{\partial U(r)}{\partial r},$$

where  $\mu$  is the reduced mass and  $U$  is an arbitrary potential. Set the radius  $r = \rho + x$ , where  $x$  is a small perturbation.

(a)

Starting from the differential equation for  $r$  and using the fact that the angular momentum  $l$  is constant, substitute  $r = \rho + x$ . Retaining terms only to first order in  $x$ , Taylor expand  $g(r)$  about the point  $r = \rho$ , and show that  $x$  satisfies the differential equation

$$\ddot{x} + \left[ \frac{3g(\rho)}{\rho} + g'(\rho) \right] x = 0 ,$$

where  $g'(\rho)$  is  $dg/dr$  evaluated at  $r = \rho$ .

(b)

Taking the force law to be  $F(r) = -kr^{-n}$ , where  $n$  is an integer, show that the angle between two successive values of  $r = r_{\max}$  (the “apsidal angle”) is  $2\pi/\sqrt{3-n}$ . Thus, if  $n > -6$ , show that in general a closed orbit will result only for the harmonic oscillator force and the inverse square law force.

6.

Consider the motion of a particle in a central force field  $F = -k/r^2 + C/r^3$ .

(a)

Show that the equation of the orbit can be put in the form

$$\frac{1}{r} = \frac{1 + \epsilon \cos \alpha \theta}{a(1 - \epsilon^2)} ,$$

which is an ellipse for  $\alpha = 1$ , but is a *precessing* ellipse for  $\alpha \neq 1$ .

(b)

The precessing motion may be described in terms of the *rate of precession of the perihelion*, where the term perihelion is used (loosely) to denote any of the turning points of the orbit. Derive an approximate expression for the rate of precession when  $\alpha$  is close to unity, in terms of the dimensionless quantity  $\eta = C/ka$ .

(c)

The ratio  $\eta$  is a measure of the strength of the perturbing inverse cube term relative to the main inverse square term of the force. Show that the rate of precession of Mercury’s perihelion ( $40''$  of

arc per century) could be accounted for *classically*, if  $\eta = 1.42 \times 10^{-7}$ . [Mercury’s period and eccentricity are 0.24y and 0.206, respectively.]

7.

A He nucleus with velocity  $v = 0.05c$  is normally incident on an Au foil that is 1 micron ( $1 \times 10^{-6}$  m) thick. What is the probability that it will scatter into the backward hemisphere, *i.e.* bounce off the foil? (Please supply a number.)

8.

Calculate the differential cross section  $d\sigma/d\Omega$  and the total cross section  $\sigma_T$  for the elastic scattering of a point particle from an impenetrable sphere; *i.e.*, the potential is given by  $U(r) = 0$ ,  $r > a$ ;  $U(r) = \infty$ ,  $r < a$ .